

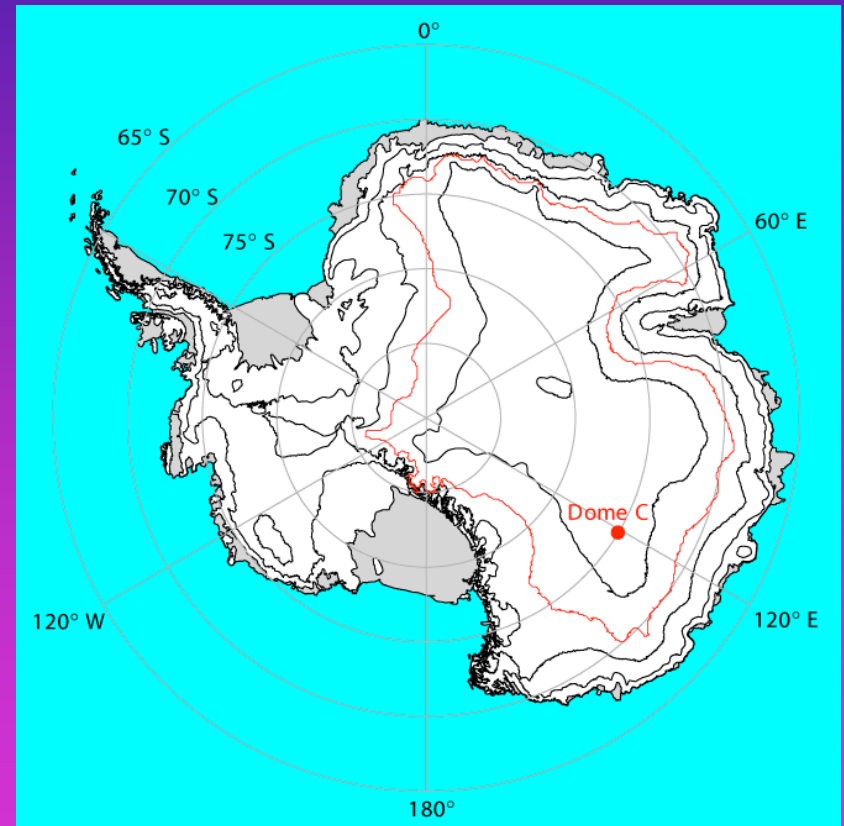
An aerial photograph of a vast, flat, snow-covered landscape, likely a tundra or a frozen body of water. The surface is covered in a dense, textured layer of snow with subtle ridges and depressions. The horizon is a straight line in the distance, and the sky is a clear, pale blue.

# A Validation of the Permanent Snow ADMs Using Data from East Antarctica

Stephen Hudson and Stephen Warren

# Observations and Parameterization

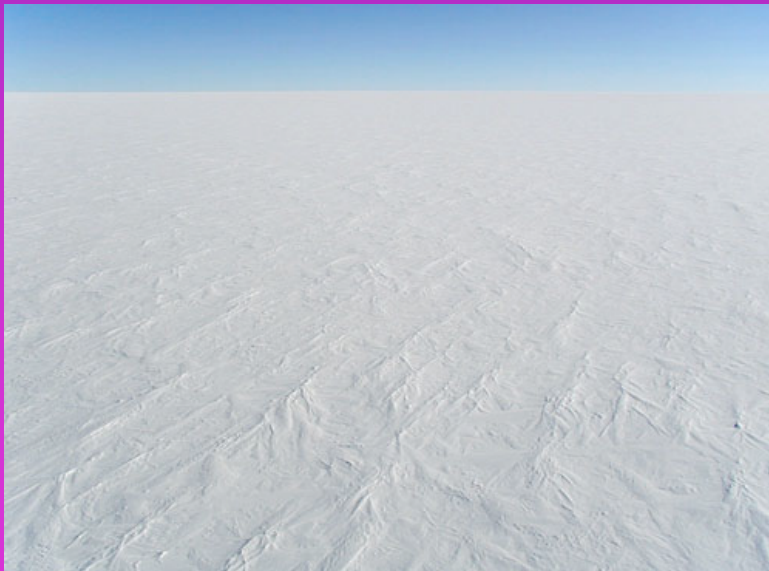
- We made spectral observations of the anisotropic reflectance factor ( $R$ ) of the snow at Dome C
  - $\lambda$  350—2400 nm
  - $\theta_o$  52—87°
- Using these, we developed a parameterization for  $R$



75°S, 123°E, 3250 m

# Observations and Parameterization

The observations were made with a  $15^\circ$  field of view from 32 m above the surface to capture the effects of the snow-surface roughness

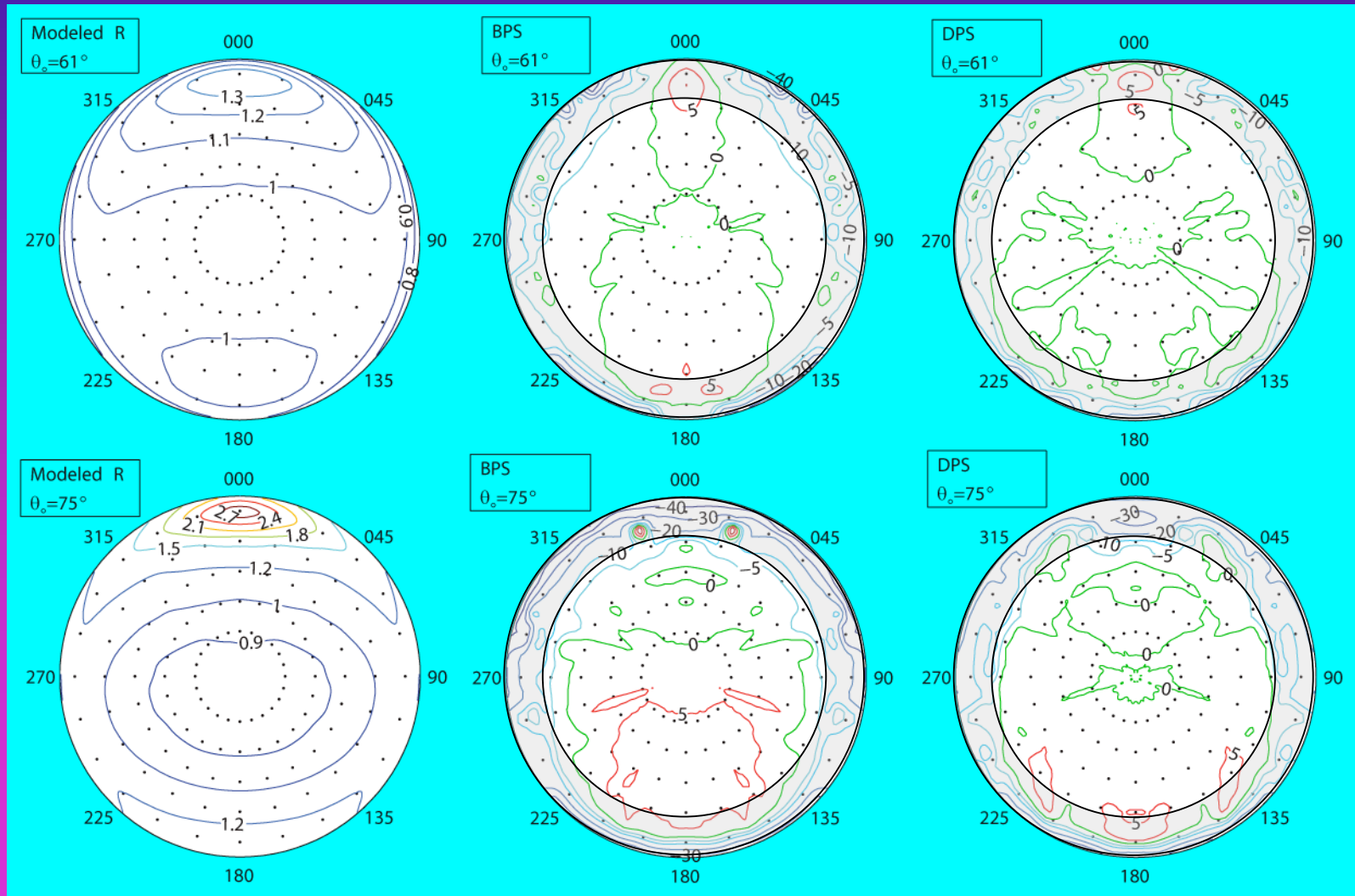




# Modelling TOA Reflectance

- The parameterization was used as the lower boundary in SBDART (DISORT), which was used to calculate TOA radiance and flux
- The atmospheric profile was the mean of 47 radiosoundings made at Dome C during January 2004, along with NOAA ozonesonde data from South Pole
- Results were integrated from 0.2 to 10  $\mu\text{m}$  for comparison with CERES shortwave channel

# Comparing CERES and Modeled ADMs

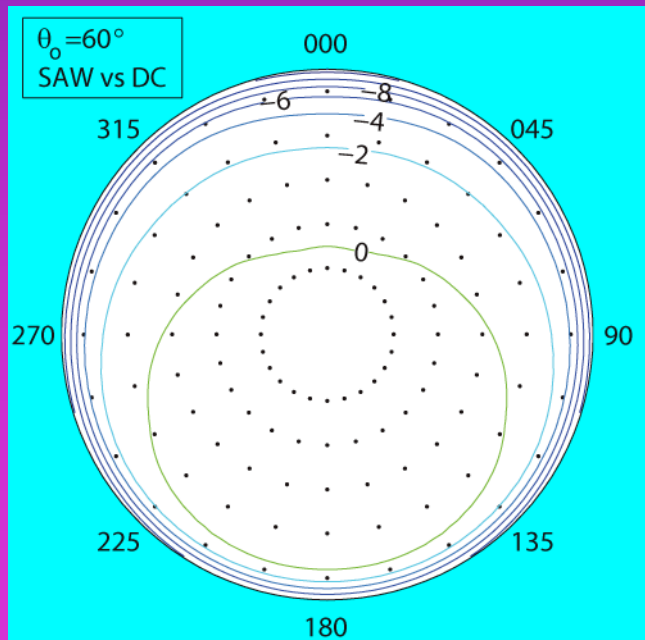


# Model

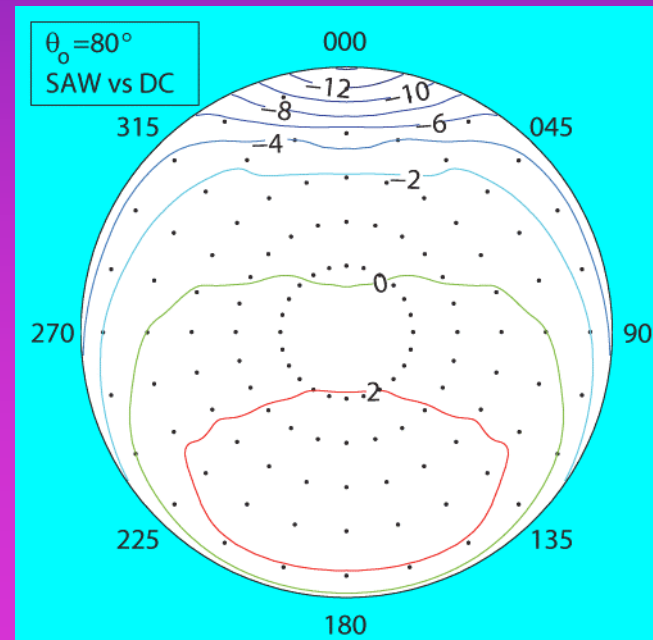
## CERES – Model (%)

# Effects of Atmospheric Variability on $R$

- The Dome-C and subarctic winter atmospheres are nearly the extreme cases for atmospheric conditions over permanent snow



$$\alpha_{DC}=0.711, \alpha_{SAW}=0.659$$

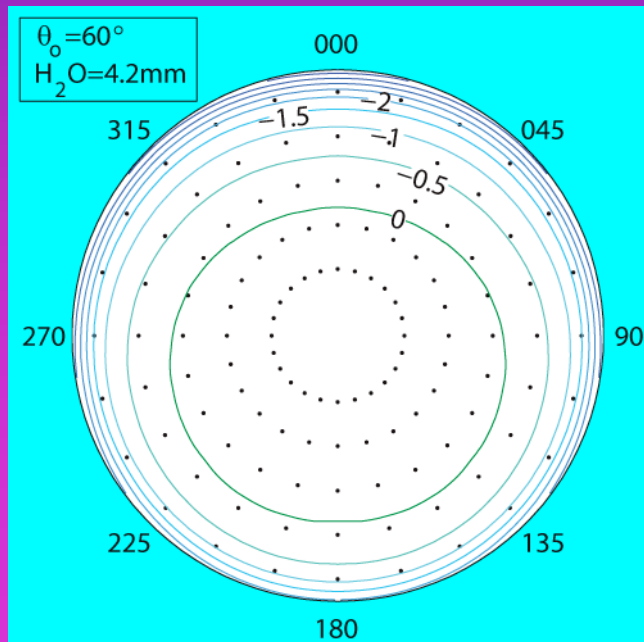


$$\alpha_{DC}=0.717, \alpha_{SAW}=0.641$$

# Effects of Atmospheric Variability on $R$

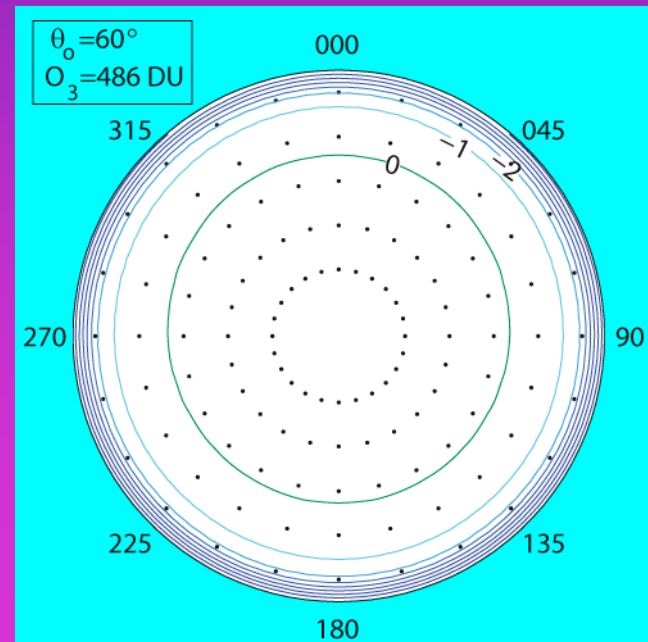
- Water vapor and ozone increases have similar effects on TOA  $R$

Increase PWV from 0.7 to 4.2 mm



$$\alpha_{0.7\text{mm}}=0.711, \alpha_{4.2\text{mm}}=0.685$$

Increase  $\text{O}_3$  from 284 to 486 DU

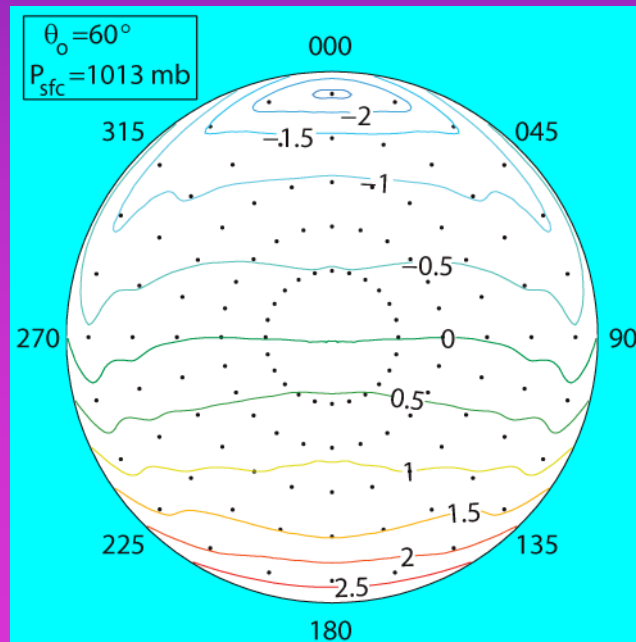


$$\alpha_{284\text{DU}}=0.711, \alpha_{486\text{DU}}=0.695$$

# Effects of Atmospheric Variability on $R$

- Surface pressure (elevation) changes also change TOA  $R$ ; change is larger, relative to albedo change

Increase  $P_{\text{sfc}}$  from 650 to 1013 mb

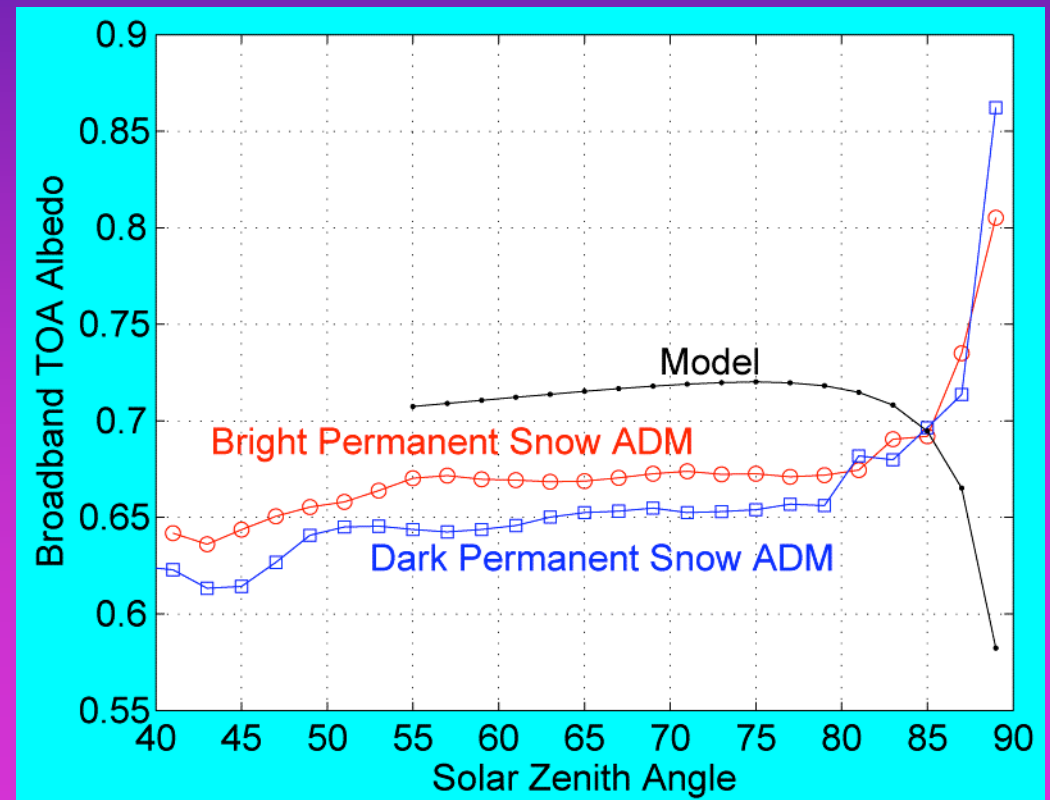


$$\alpha_{650\text{mb}} = 0.711, \alpha_{1013\text{mb}} = 0.702$$



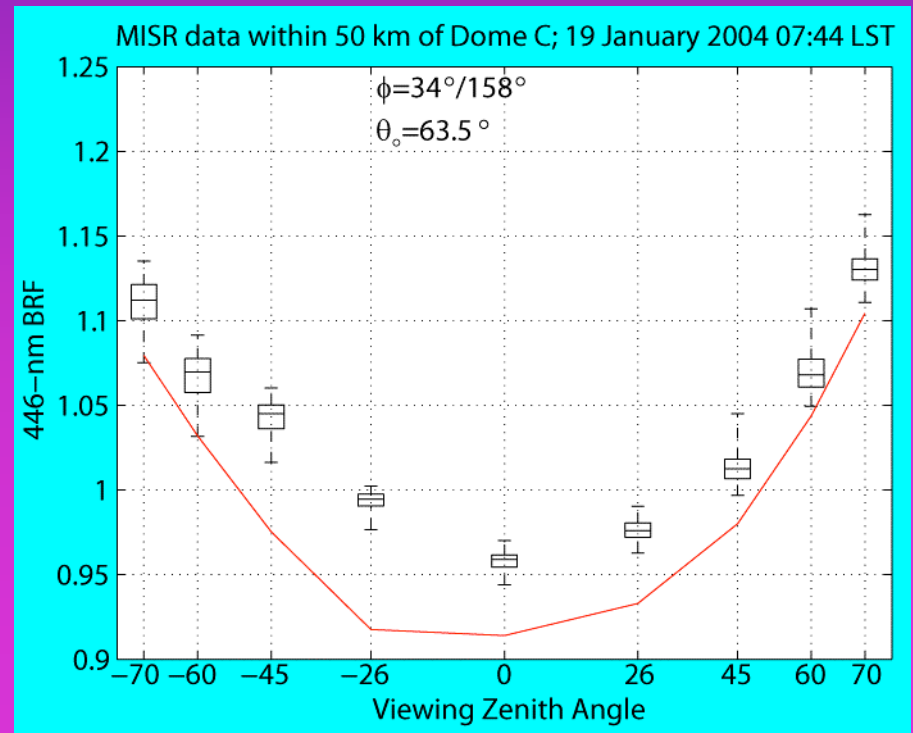
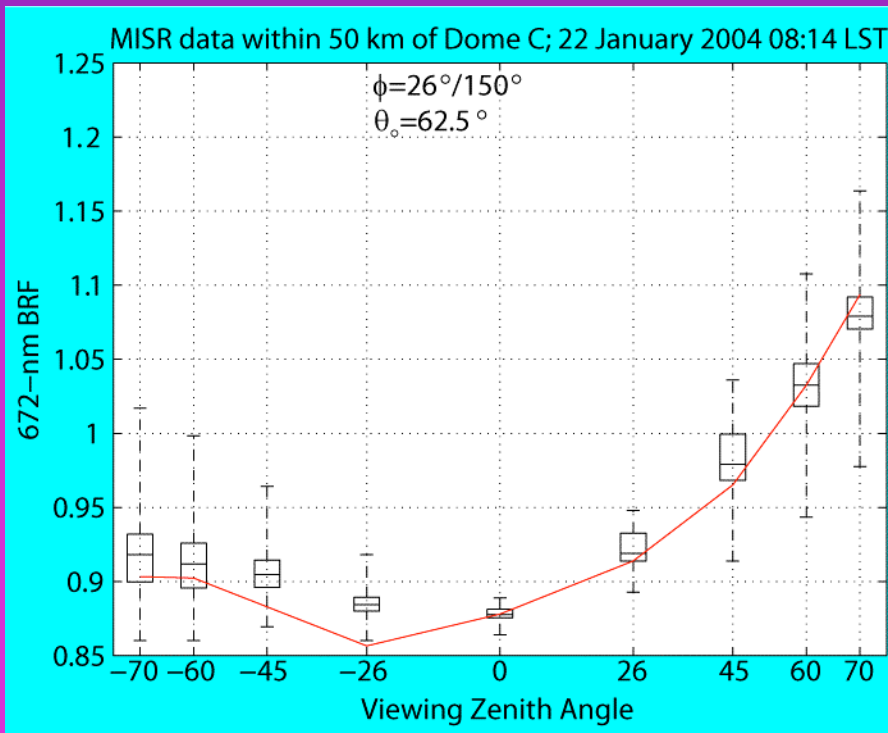
# CERES and Modeled Albedo Comparison

- Modeled albedos at TOA are larger by nearly 10% than the ADM albedos
- Some of this difference is expected since Dome C is higher and drier than many permanent snow regions



# Is the Model Overestimating Reflectance?

- Comparisons of the modeled radiances with MISR observation suggest the model gives accurate or low estimates of radiance.



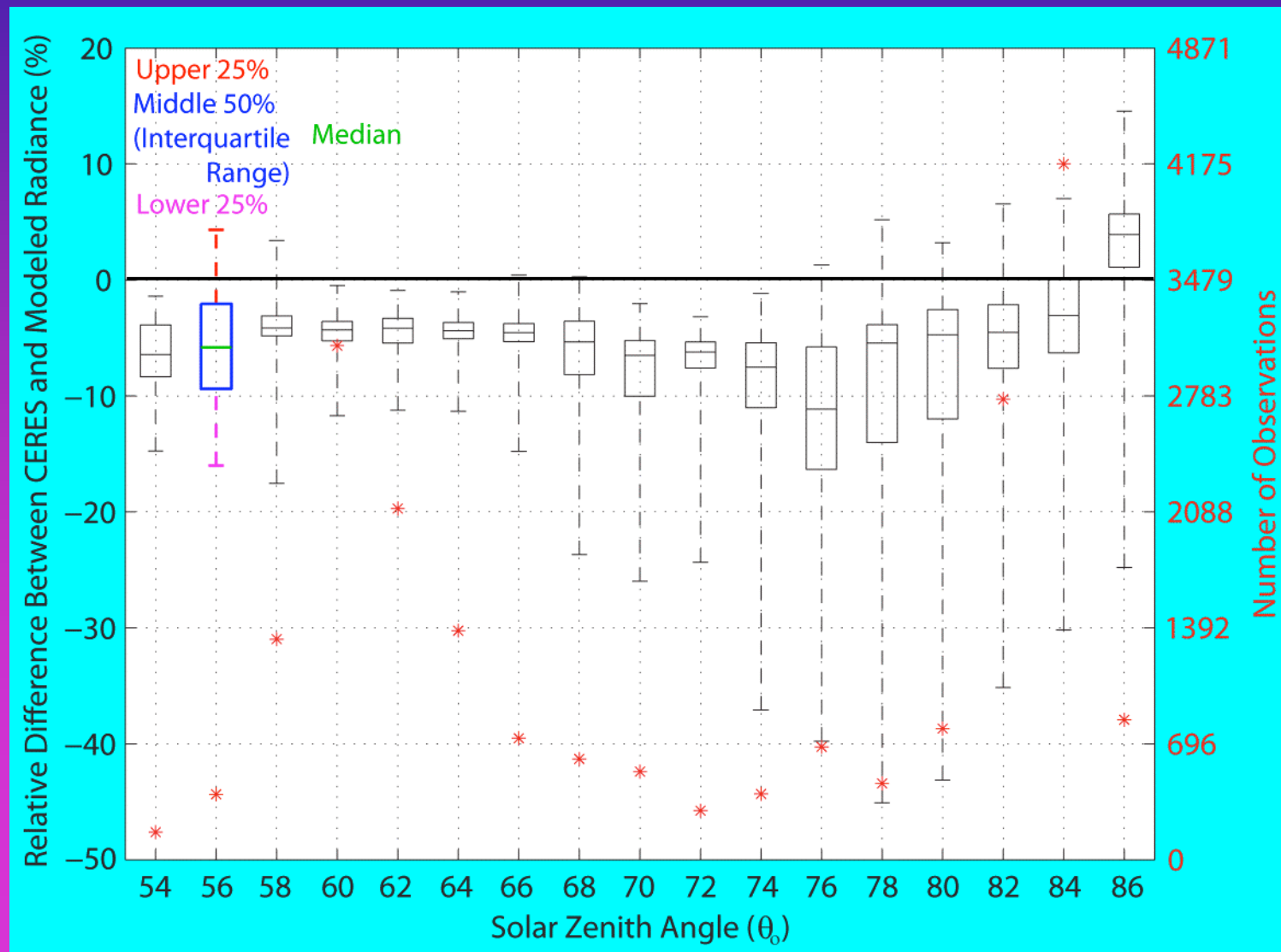
# Conclusions

- The CERES permanent snow ADMs are representative of the Dome-C region
- Atmospheric variability over permanent snow regions can cause variations in R at the TOA of generally less than 2% at viewing zenith angles less than 70°

# Conclusions

- The average albedo of permanent snow from the ADMs is too low for the Dome-C region
- Some of the albedo difference is due to the thin, dry atmosphere above Dome C
- Direct comparisons of radiance observations with the model show about a 4% bias, with CERES radiances less than modeled

# Assessing CERES SW calibration





## Possible explanations for CERES-Model radiance offset

### A. Could measured surface albedo be unrepresentative; higher than in CERES footprint?

1. The albedo was measured over a flat surface; albedo of rough surface is lower.

*Insignificant at Dome C (albedo difference 0.003)*

*Bigger difference where sastrugi are larger, but blowing snow hides sastrugi*

2. Soot contamination near the station?

*Error is in the wrong direction. Soot at station was 1- 7 ppb;*

*in remote snow 0.2-0.7 ppb*

3. Grain size smaller at measurement site?

*Would need  $r=300\text{ }\mu\text{m}$*

*Traverse showed no variation; grain radius 50-100  $\mu\text{m}$  along route.*

## B. Could measured surface albedo be erroneously high?

### 1. Shadowing correction too large?

*Correction of 4% was obtained in two different ways.*

*If correction were removed, albedo would be too low in visible by 0.04  
based on known absorption coefficient of ice.*

*Correction was 1% at South Pole; gave same visible albedo.*

### 2. Cosine-collector inaccuracy.

*Measurement was made under diffuse illumination; effect cancels.*

*Multi-year albedos from pyranometer at South Pole have been corrected  
for deviations from "cosinicity"*

### 3. Surface slope causes errors.

*No error under diffuse light.*

*Dome-C region has slope  $< 0.06$  degrees*

## C. Modeling errors

### 1. Dependence of albedo on incidence angle

*Discrepancy with CERES is independent of solar zenith angle.  
Modeled surface albedo is actually too low for blue and UV  
because of old ice optical constants.*

### 2. Atmospheric aerosols

*Adding stratospheric aerosol to model slightly **increases**  
computed TOA albedo.*

### 3. Water-vapor and ozone column amounts

*Multiplying water-vapor by 6 would decrease TOA albedo by 3.6%.  
Doubling ozone would decrease TOA albedo by 3%.*